

(Dated: February 15, 2010)

The numerical simulations were performed using VORPAL, a conformal finite-difference time-domain (CFDTD) particle-in-cell electromagnetic solver [9]. VORPAL is a parallel, object-oriented framework for three dimensional relativistic electrostatic and electromagnetic plasma simulation. We extensively benchmarked our initial VORPAL simulations of the SPFEL process against earlier work [10]; see Ref. [11]. The geometric configuration of our simulation model is shown in Fig. 1. The model includes perfectly-conducting rectangular gratings, a particle source on the left end side and a “particle sink” on the right that allows macroparticles to exit the computa-

TABLE I: Grating and beam parameters used for the VORPAL simulations.

Parameter	Symbol	Value	Unit
Grating Period	λ_g	200	μm
Groove Width	w	100	μm
Groove Depth	h	100	μm
Number of Periods/grating	N_g	75	—
Electron Energy	E	25-200	keV
Beam Current	I	≥ 135	A/m
Beam Thickness	b	20-400	μm
Beam Clearance	g	20-30	μm
Grating gap	$2x_0$	$b + 2g$	μm
External Magnetic Field	B_a	2.0	T

tional domain without being scattered or creating other source of radiation. The grating grooves are along the \hat{z} direction and the beam propagates along the \hat{y} axis. Beside the grating, the other surrounding boundaries consist of perfectly matched layers that significantly suppresses artificial reflections of incident radiation. The particle source produces a uniformly-distributed (in all directions) DC beam with an instantaneous risetime and is transversely confined by an uniform external axial magnetic field $\mathbf{B}_a = B_a \hat{y}$. The simulations parameters used for the simulations presented below are gathered in Tab. I, with $E = 50$ keV, $I = 135$ A.m $^{-1}$, $b = 50$ μm , and $g = 20$ μm ; for the sake of simplicity the buncher and radiator gratings are identical.

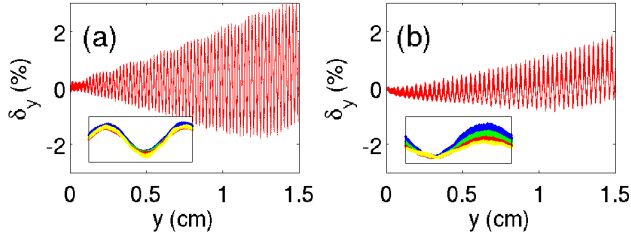


FIG. 2: (Color online) Snapshot of the axial phase spaces (y, δ_y) (where $\delta_y = \gamma\beta_y/[\beta_y(y=0)\gamma(y=0)] - 1$ is the fractional axial momentum spread) for a double (a) and single (b) grating taken at 151 ps. The lower insets on each of the plot present zoomed-in phase space over for $y \in [0.50, 0.53]$ cm, for which the macroparticles have been color-coded accordingly to their height within the beam. The double-grating simulation has less distinct layers, signifying more uniform bunching across the height of the electron beam. The beam parameters are $E = 50$ keV and $I = 135$ A.m $^{-1}$.

We first analyze the buncher section of the proposed two-stages SPFEL. Contrary to the SPFEL configurations explored in previous work, the beam propagates between two identical gratings arranged symmetrically; see Fig. 1. Considering the two-dimensional problem, the evanescent waves produced from the beam passing over each grating located at $\pm x_0$ are $E_{y,e}^{\pm}(x, y) = E_{0,e}(y)e^{-\alpha(x_0 \mp x)}$ (where \pm respectively designate the

electric field contributions from the upper and lower grating). The overlap of the evanescent waves from both gratings results in an evanescent field $E_{y,e}(x, y) = 2E_{0,e}(y)e^{-\alpha x_0} \cosh(\alpha x)$. Beside being stronger by a factor of two (for $x = 0$) thereby bunching the beam stronger, the field is also more uniform across the beam's thickness and effectively result in a more uniform bunching. The strengthened evanescent field may also be taken advantage of to relax the start current requirements. These features are confirmed by numerical simulation; see Fig. 2. The velocity modulation is approximately twice as large for the double grating configuration than for a single grating configuration. The presence of the second grating also affects the dispersion relation and λ_e , which in-turn leads to a different bunching frequency than the single grating system; see analysis in Ref. [12]. This feature offers a greater flexibility for tuning the evanescent wave frequency (and therefore the radiation frequency) by either varying the electron beam energy or altering the gap between the two gratings as demonstrated in Fig. 3. As the gap increases, the frequency converges with that of a single-grating system. Over the considered range of energy $E \in [25, 200]$ keV and grating gap $2x_0 \in [90, 440]$, significant velocity modulation is observed and the corresponding evanescent wavelength variation of respectively $\sim 15\%$ and $\sim 35\%$ are approximately two orders of magnitude than the produced radiation spectrum width as will be shown later.

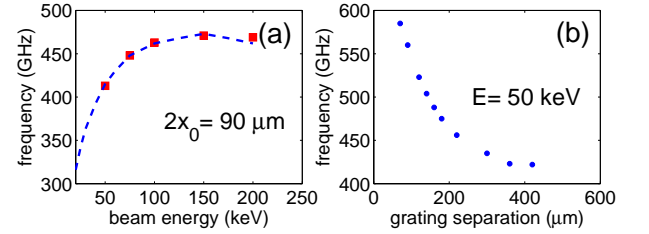


FIG. 3: (Color online) Evanescent wave frequency versus beam energy for the single-grating configuration (a) and evanescent wave frequency versus grating separation for a double-grating configuration (b). The markers are results from VORPAL simulations while the dashed line is obtained from the dispersion relation derived in Ref. [2].

The second stage of the proposed SPFEL consists of passing the microbunched beam close to a single grating. The spectral intensity radiated by a bunch of N electrons via the SP effect is related to the single electron intensity $\frac{dI}{d\omega}|_1$ via $\frac{dI}{d\omega}|_N = \frac{dI}{d\omega}|_1 [N + N^2|S(\omega)|^2]$ where $S(\omega)$ is the intensity-normalized Fourier transform of the charge distribution $S(t)$ [13]. Considering a series of N_b identical microbunches with distribution $\Lambda(t)$ we have $S(t) = \sum_{n=1}^{N_b} \Lambda(t + nT)$ (where $T = \lambda_e/c$ is the period) giving $|S(\omega)|^2 = \xi |\Lambda(\omega)|^2$. The intra-bunch coherence factor $\xi \equiv \sin^2(\omega N_b T/2) / [\sin^2(\omega T/2)]$ describes the enhancement of radiation emission at resonance, i.e. for frequencies $\omega = 2\pi n c / \lambda_e$.

Figure 4 shows a contour plot of the E_y -field over

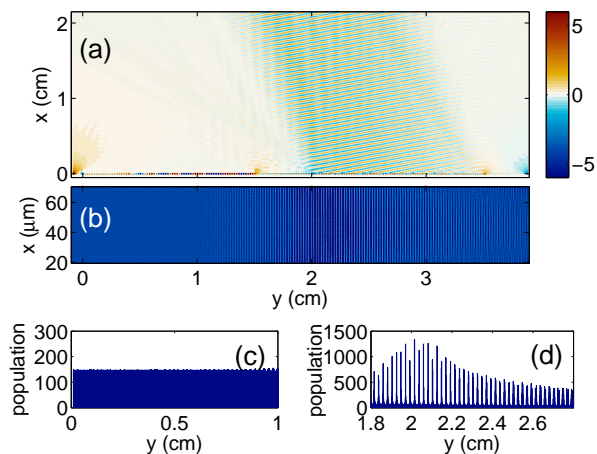


FIG. 4: (Color online) Snapshots at $t = 906$ ps of the axial electric field E_y and corresponding spatial distribution of the beam in the $(x, y, z = 0)$ plane for a two-stage SPFEL (the colorbar units are $\text{MV} \cdot \text{cm}^{-1}$). The associated axial projections at two axial locations are shown in (c) and (d).

the computational domain and clearly demonstrates the emission of SP radiation with a planar wavefront from the second grating. The planar wavefront is a signature of superradiance due to constructive interference between successive bunches. The wavevector associated to this radiation makes a 106° angle and the wavelength value $\simeq 540 \mu\text{m}$ is consistent with the one expected from the SP relation [5] $\lambda_m = \lambda_g / |m|(\beta^{-1} - \cos \theta) = 540 \mu\text{m}$ (taking $m = 1$, $\lambda_g = 200 \mu\text{m}$ and $\beta = 0.4127$). Whereas the single-grating SPFEL requires that the super-radiant emission occurs at $m \geq 2$, the proposed two-stage SPFEL can operate in the super-radiant mode at $m = 1$ by a proper selection of parameters. Even if the radiator and the buncher have the same grating period, the increased bunching frequency from the double-grating dispersion relation may match that of the $m=1$ mode of the SP relation. Operating at $m = 1$ enables the generation of higher radiation rates compared to $m \geq 2$. Figure 4 also displays a snapshot of the spatial distribution and associated projections confirming the formation of microbunches.

The time evolution of the magnetic field B_z recorded

at $\theta = 106^\circ$ and $R = 2$ cm along with its spectrum are presented in Fig. 5. The spectrum shows that significant emission occurs at $f_e = 556$ GHz, the frequency of the evanescent wave supported by the buncher section, and its harmonics. The bandwidth of the radiation emitted at the first harmonic is $\delta f \simeq 1.3$ GHz resulting in $df/f_e \simeq 0.23\%$. The spectrum also displays a 413 GHz frequency component which corresponds to the evanescent wave produced by the downstream grating. Taking the wavefront to be a plane wave with steady-state peak magnetic field of $B_0 \sim 40 \mu\text{T}$ (see Fig. 5), we estimate the time-averaged Poynting vector to be $\langle S \rangle = cB_0^2/(2\mu_0) \simeq 2 \times 10^4 \text{ W} \cdot \text{m}^{-2}$. Harvesting the radiation using a $1 \times 1 \text{ cm}^2$ mirror would result in a total average power of $P \sim 2 \text{ W}$.

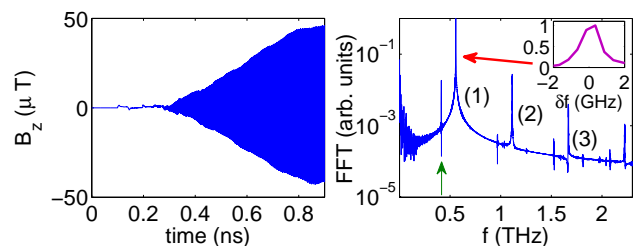


FIG. 5: (Color online) Time evolution of the magnetic field B_z at $(R, \theta) = (2 \text{ cm}, 106^\circ)$ (a) and associated FFT (b). The beam parameters are $I = 135 \text{ A} \cdot \text{m}^{-1}$ and $E = 50 \text{ keV}$. The inset in (b) is a zoom-in of the 1st harmonic peak showing the bandwidth of 1.3 GHz (FWHM). The numbers (1), (2) and (3) indicate the harmonics of the evanescent wave associated to the double grating configuration. The green vertical arrow indicates the evanescent wave supported by the downstream grating.

In summary we have demonstrated several advantages of a two-stage SPFEL and show that the device can be used to produce tunable superradiant THz radiation with average power of the order of Watts.

We thank Dr. P. Spentzouris (FNAL) for granting us access to the NERSC computers where our simulations were performed. This work was partially supported by the US Department of Education under contract P116Z050086 with Northern Illinois University.

-
- [1] M. Tonouchi, *Nature Photonics*, **1**, 97-105 (2007).
 - [2] H. L. Andrews and C. A. Brau, *Phys. Rev. ST Accel. Beams*, **7**, 070701 (2004).
 - [3] V. Kumar and K.-J. Kim, *Phys. Rev. E* **73**, 026501 (2006).
 - [4] J. M. Wachtel, *J. Appl. Phys.*, **50** (1) (1979).
 - [5] S. J. Smith and E. M. Purcell, *Phys. Rev. Lett.* **92**, 1069 (1953).
 - [6] H. L. Andrews, *et al.*, *Proceedings of the 2004 FEL conference*, 278 (2004).
 - [7] D. Li, *et al.*, *Appl. Phys. Lett.* **88**, 201501 (2006).
 - [8] Y. Li and K.-J. Kim, *Appl. Phys. Lett.* **92**, 014101 (2008).
 - [9] C. Nieter and J. R. Cary, *J. Comp. Phys.*, **196**, 448 (2004).
 - [10] J. T. Donohue and J. Gardelle, *Phys. Rev. ST Accel. Beams* **8**, 060702 (2005).
 - [11] C. R. Prokop, P. Piot, M. C. Lin, and P. Stoltz, *Proceedings of the 2009 FEL conference*, 180 (2009).
 - [12] H. P. Freund and T. M. Abu-Elfadl, *IEEE Trans. Plasma Sci.*, **32**, 1015 (2004).
 - [13] J. S. Nodvick and D. S. Saxon, *Phys. Rev.* **96**, 180 (1954).